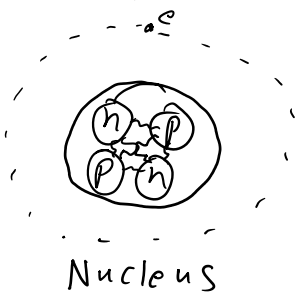
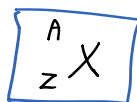


Composition of the Nucleus



for an element X:



A → Mass number = No. of protons + Neutrons
Z → Atomic number

Size of the Nucleus

Radius of a nucleus: $R = R_0 A^{1/3}$

$$R_0 = 1.2 \times 10^{-15} \text{ m} = 1.2 \text{ fm}$$

$$10^{-15} \text{ m} = \text{fm (fermi)}$$

Volume of the nucleus $\propto R^3 \propto A$

$$V = \frac{4}{3} \pi R^3 \propto R^3 \propto (A^{1/3})^3 \propto A$$

* Density $\propto \frac{\text{Mass}}{\text{Volume}} \propto \frac{A}{A} = \text{Constant}$

→ Derivation

Nuclei density is constant

$$\text{Nuclear density} = 2.3 \times 10^{17} \text{ kg m}^{-3}$$

Example 13.1 Given the mass of iron nucleus as 55.85 u and $A = 56$, find its nuclear density. (Given $1 \text{ u} = 1.66 \times 10^{-27} \text{ kg}$)

MASS-ENERGY EQUIVALENCE

$$E = mc^2$$

\downarrow \downarrow \downarrow
 J kg ms^{-1}

$$c = 3 \times 10^8 \text{ m/s}$$

↳ Speed of light in vacuum.

$$1 \text{ u} \equiv 931.5 \text{ MeV}$$

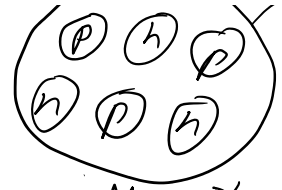
Eq 13.2 :- Calculate the energy equivalent of 1 g of substance.

* Nuclear Binding Energy

$$\text{Mass defect: } \Delta M = [Zm_p + (A-Z)m_n] - M$$



Mass defect: $\Delta M = [Zm_p + (A-Z)m_n] - M$



$M < 3(m_p + m_n)$

Binding Energy: $E_b = \Delta M c^2$ (Energy Released)

Ex 13.3 Find the energy equivalent of one atomic mass unit, first in Joules, then in MeV.

Using this, express the mass defect of ${}^8_8\text{O}$ in MeV/c^2 .
 (Given: $m_p = 1.00727u$, $m_n = 1.00866u$, $m_0 = 15.99053u$)
 ($1u = 1.660539 \times 10^{-27} \text{kg}$)

$\Delta m = (8m_p + 8m_n) - m_0 = X u = 931.5 \times X \text{ MeV}/c^2$

Sol:- $E = mc^2$

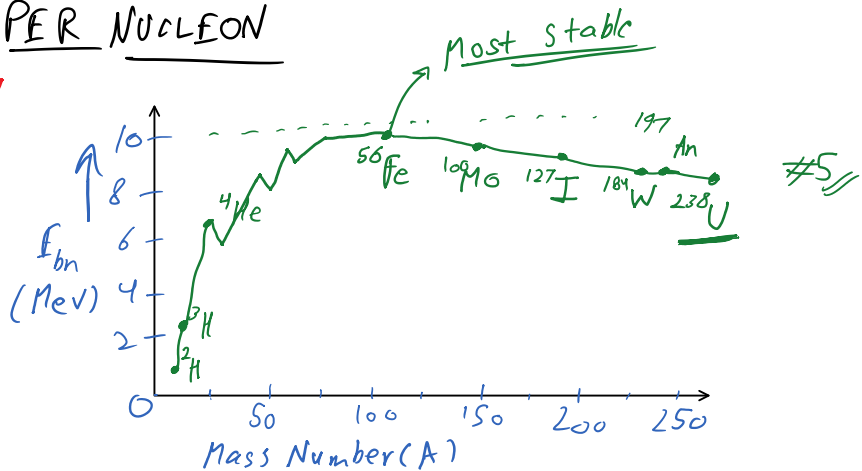
$E = (1u)(3 \times 10^8)^2$

$= \frac{(1.660539 \times 10^{-27})(3 \times 10^8)^2}{1.6 \times 10^{-19} \text{ eV}} = 931.5 \text{ MeV}$

$\Delta m = [Zm_p + (A-Z)m_n] - N = [8m_p + (16-8)m_n] - m_0 \Rightarrow \frac{\Delta m c^2 \text{ MeV}}{c^2} = \phi \text{ MeV}/c^2$

BINDING ENERGY PER NUCLEON

$E_{bn} = \frac{E_b}{A}$



NUCLEAR FORCE
(Strong)

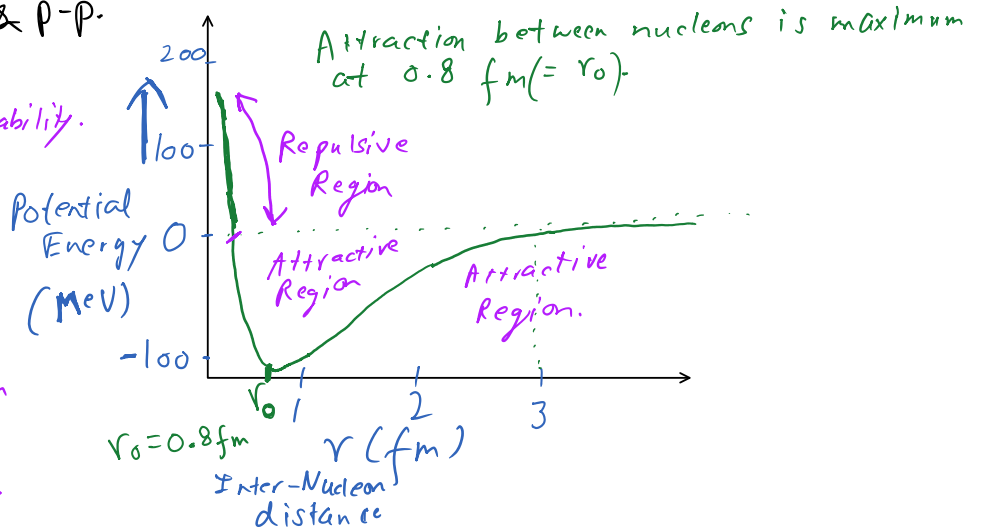
(i) Nuclear Force is much stronger than the Coulomb force or gravitational force.

- (ii) Nuclear force between two nucleons rapidly falls to zero as their distance becomes more than a few femtometers.
- (iii) Nuclear force does not depend on charge. It is approximately same between n-n, p-n & p-p.

▷ Less energy means more stability.

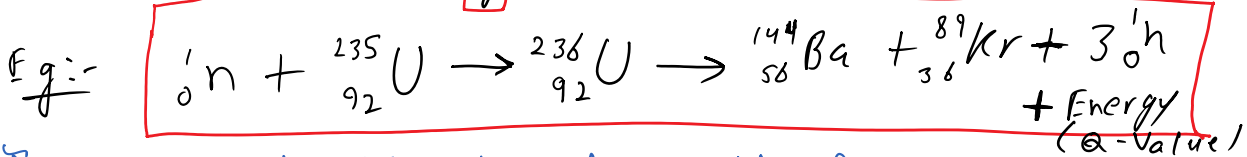
▷ Nuclear force is attractive when P.E. < 0.

▷ The minima of P.E. in the graph is the region of most stability & strongest attraction.



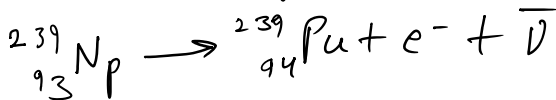
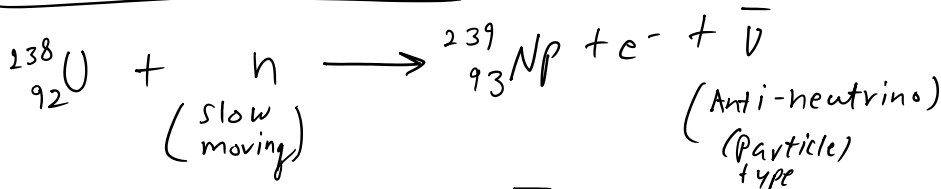
NUCLEAR FISSION

* Larger Nucleus breaking into smaller Nuclei.

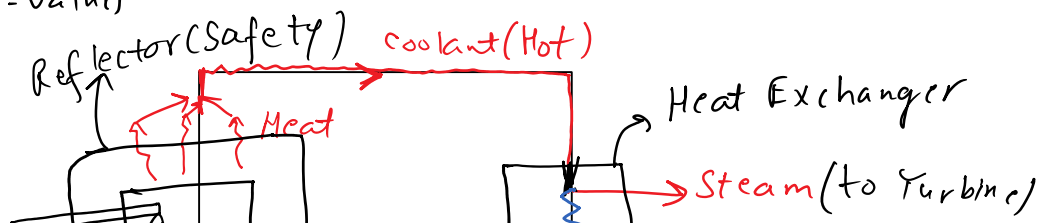


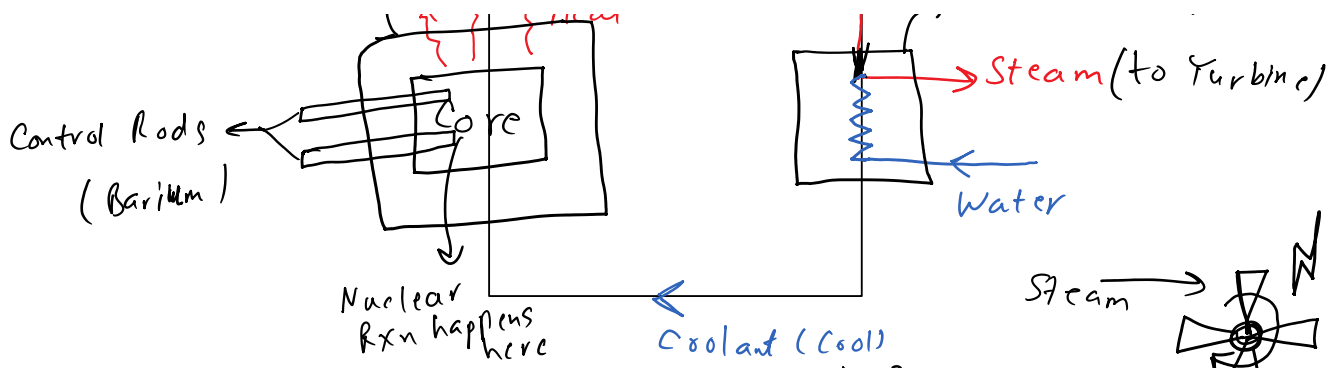
* The energy released (Q-Value) in the fission reaction of nuclei like Uranium is of the order of 200 MeV per fissioning nucleus.
 ↳ When 1 ${}_{92}^{236}\text{U}$ breaks, it gives out around 200 MeV of Energy (Heat)

Nuclear Reactor



* Then, Plutonium undergoes fission with slow neutrons to give nuclear Energy (mostly Heat) (Q-Value)



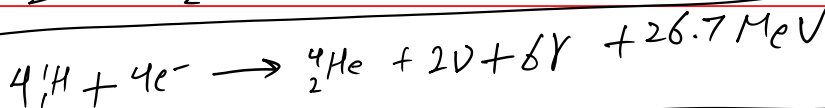
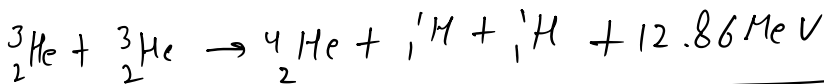
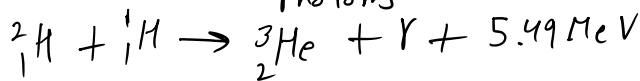
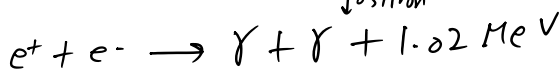
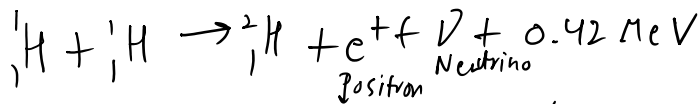


H.W. Exercise

$$\begin{array}{r} 13.1 - 13.5 \\ \hline 13.15 - 13.17 \\ \hline \end{array} \quad \begin{array}{r} 13.20 \\ \hline 13.21 \\ \hline \end{array}$$

Nuclear Fusion

↳ Small nuclei combining to give larger nuclei/nucleus.

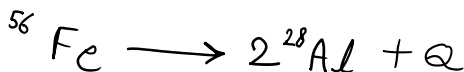


13.1 (a) $\frac{75}{100} M_{6,3} + \frac{25}{100} M_{7,3}$ (b) $\frac{x}{100} M_{10,5} + \frac{(100-x)}{100} M_{11,5}$

13.2

13.16 $Q = \Delta mc^2 = (M_p - M_r) c^2 = (2M_{Al} - M_{Fe}) c^2$

If $Q > 0$, then energetically feasible
If $Q < 0$, " " " not " "



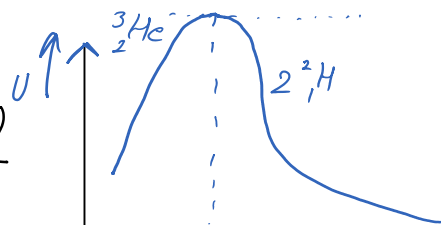
13.17 $M_{Pu} \approx 239 \text{ u}$ $Q = 180 \text{ MeV}$

$$E = \frac{1 \text{ kg}}{239 \times 1.66 \times 10^{-27}} \times 180 \text{ MeV}$$

13.20

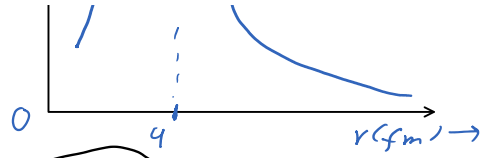


$$U = \frac{1}{4\pi\epsilon_0} \frac{(+e)(+e)}{(4 \times 10^{-15})}$$





$$V = \frac{1}{4\pi\epsilon_0} \frac{1}{(4 \times 10^{-15})}$$



$$T = \frac{1}{f} \therefore f = \frac{1}{T}$$

$$T = \frac{2\pi r v}{v} \quad \left(v = r_0 \frac{h^2}{Z} \right)$$

53 pm
0.53 Å

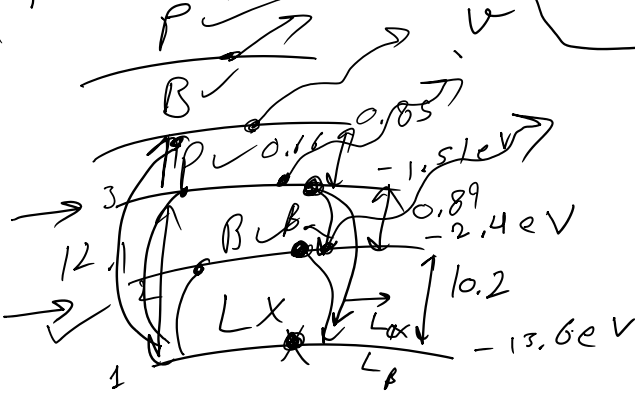
$$v = v_0 \frac{Z}{n}$$

$$v_0 = 2.2 \times 10^8$$

9

$$\frac{151}{-85}$$

$$066$$



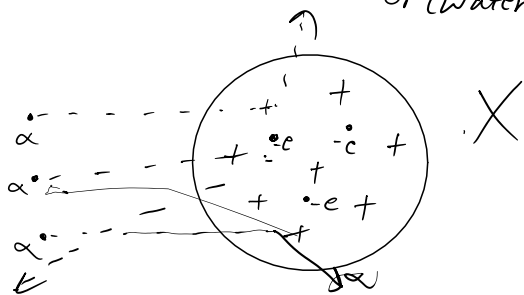
$$\frac{E_0}{n^2} = \frac{13.6}{1}$$

$$\frac{13.6}{16}$$

$$\frac{13.6}{4}$$

$$\frac{13.6}{9}$$

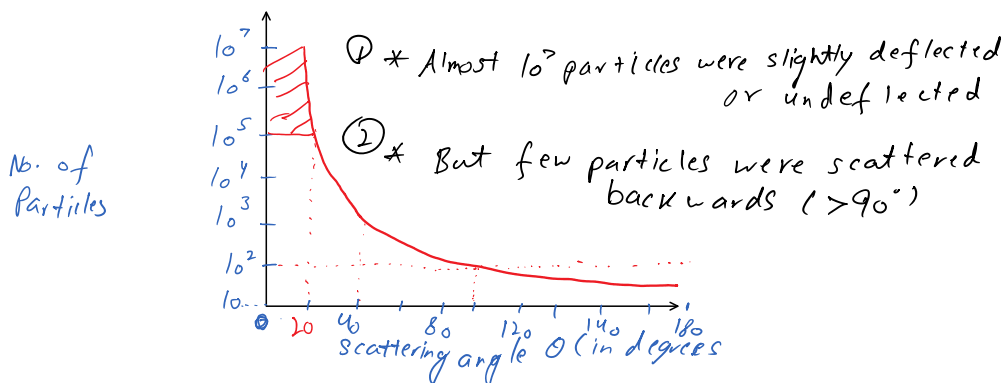
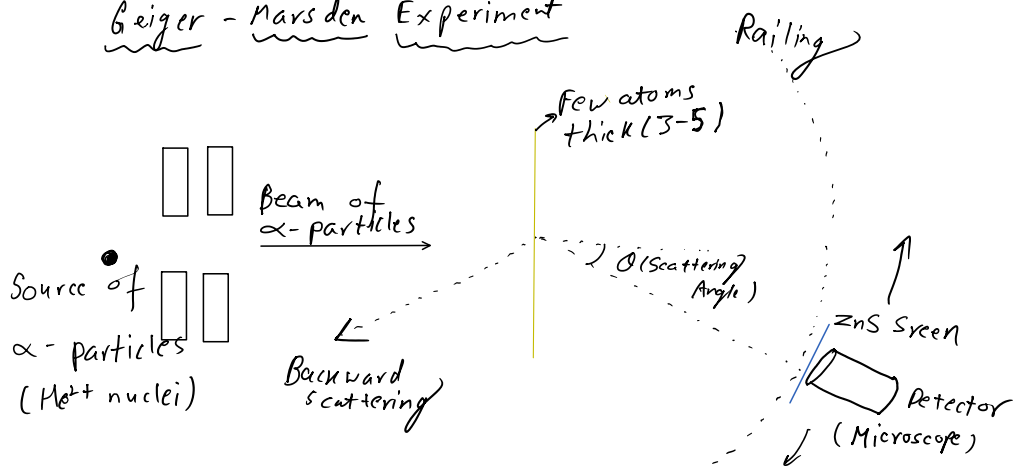
Thomson's Model of Atom (Plum Pudding Model)
or (Watermelon Model)



J.J. Thomson discovered the electron.
James Chadwick " " neutron.

Rutherford's Alpha-Scattering Experiment

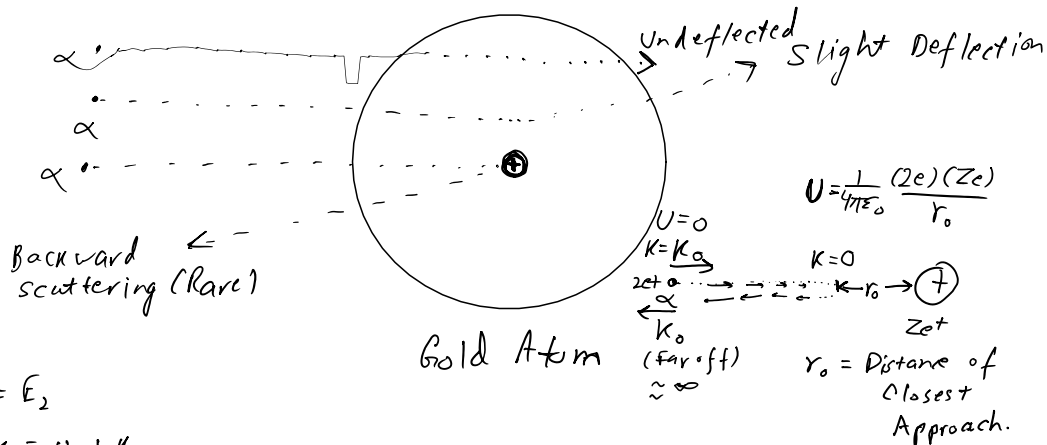
Geiger - Marsden Experiment



① Most of the atom is empty space.

② Most of the mass & all positive charge resides in a small region at the centre of the atom called the Nucleus

Hence, E. Rutherford DISCOVERED THE NUCLEUS



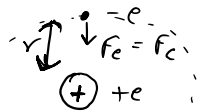
$$U_1 + K_1 = U_2 + K_2$$

$$0 + K_0 = \frac{1}{4\pi\epsilon_0} \frac{(2e)(2e)}{r_0} + 0$$

ELECTRON ORBITS (H-atom)

$r \rightarrow$ Radius of H-atom
Bohr radius

$$\approx 53 \text{ pm or } 5.3 \times 10^{-11} \text{ m}$$



Electrostatic Force = Centripetal Force

$$\frac{1}{4\pi\epsilon_0} \frac{(e)(e)}{r^2} = \frac{mv^2}{r}$$

$$\Rightarrow v = \frac{e^2}{4\pi\epsilon_0 m r^2}$$

$$K = \frac{1}{2} mv^2 ; U = \frac{1}{4\pi\epsilon_0} \frac{(+e)(e)}{r}$$

$$\text{Total Energy, } E = K + U = \frac{e^2}{8\pi\epsilon_0 r} - \frac{e^2}{4\pi\epsilon_0 r}$$

$$E = -\frac{e^2}{8\pi\epsilon_0 r}$$

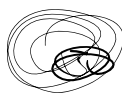
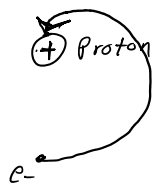
$$E = -K = \frac{U}{2} \quad \text{Imp}$$

Eq 12.3 :- $E = -13.6 \text{ eV}$

H.W. Ex - 1, 2

DRAWBACKS OF RUTHERFORD'S MODEL OF ATOM

- * According to Electromagnetic theory, an accelerated charged particle should emit Electromagnetic Radiation & Hence lose Energy.
- * By this logic, an electron in an atomic orbit should lose energy & hence collapse into the nucleus.
- * Therefore, matter cannot be stable according to Rutherford's model.
- * This dilemma was solved by Niels Bohr in his atomic model.



BOHR'S MODEL OF ATOM

(i) Bohr's first postulate was that an electron in an atom could revolve in certain stable orbits without the emission of radiant energy, contrary to the electromagnetic theory. These stable orbits are called stationary states of the atom.

(ii) Bohr's second postulate defines these stable orbits.

$$L = \frac{nh}{2\pi}$$

$L \rightarrow$ Angular momentum $= mvr$
 $n \rightarrow$ whole number
 $h \rightarrow$ Planck's constant
 $h = 6.6 \times 10^{-34} \text{ Js}$

(iii) An electron might make a transition from one stationary state to another by emitting or absorbing a photon. The energy of photon is equal to the energy difference

between the two states.

$$h\nu = |E_f - E_i|$$

$$* L_n = m v_n r_n = \frac{n h}{2\pi}$$

$$\text{But, } v_n = \frac{e}{\sqrt{4\pi\epsilon_0 m r_n}}$$

$$\therefore v_n = \frac{1}{n} \frac{e^2}{4\pi\epsilon_0} \frac{1}{h/2\pi}$$

$$\therefore r_n = \left(\frac{n^2}{m}\right) \left(\frac{h}{2\pi}\right)^2 \frac{4\pi\epsilon_0}{e^2}$$

$$\therefore E_n = \frac{-m e^4}{8 n^2 \epsilon_0^2 h^2}$$

$$\Rightarrow v_n \propto \frac{Z}{n}$$

$$\Rightarrow r_n \propto \frac{n^2}{Z}$$

$$\Rightarrow E_n \propto \frac{Z^2}{n^2}$$

Z = Atomic Number
n = Orbital Number

V. Imp

$$\Rightarrow v_n = \frac{Z}{n} v_0$$

$$v_0 = 2.2 \times 10^6 \text{ m/s}$$

$$\Rightarrow r_n = \frac{n^2}{Z} r_0$$

$$r_0 = 5.3 \times 10^{-11} \text{ m}$$

$$\Rightarrow E_n = \frac{Z^2}{n^2} E_0$$

$$E_0 = -13.6 \text{ eV}$$

▷ ENERGY LEVELS

▷ ENERGY SPECTRA

HW Revise Notes

ENERGY LEVELS

(H-atom) $\Rightarrow Z=1$

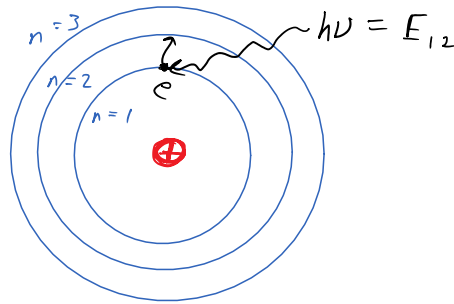
$$E_n = -13.6 \text{ eV} \left(\frac{Z^2}{n^2}\right) = \frac{-13.6 \text{ eV}}{n^2}$$

$$E_1 = -13.6 \text{ eV}$$

$$E_2 = \frac{-13.6 \text{ eV}}{2^2}$$

$$E_3 = \frac{-13.6 \text{ eV}}{3^2}$$

$$E_n = \frac{-13.6 \text{ eV}}{n^2}$$



$$E_{12} = E_2 - E_1 = -13.6 \text{ eV} \left(\frac{1}{4} - 1\right) = 10.2 \text{ eV}$$

$$\therefore h\nu \text{ or } \frac{hc}{\lambda} = 10.2 \text{ eV}$$

$$E_{ij} = E_j - E_i = \frac{hc}{\lambda}$$

$$\Rightarrow \frac{1}{\lambda} = \frac{E_j - E_i}{hc} = \left(\frac{13.6 \text{ eV}}{hc}\right) \left[\frac{1}{i^2} - \frac{1}{j^2}\right]$$

where $i < j$

$$\frac{1}{\lambda} = R \left[\frac{1}{i^2} - \frac{1}{j^2}\right]$$

metres

R \rightarrow Rydberg Constant

$$R = 1.097 \times 10^7 \text{ m}^{-1}$$

Eg:- Using the Rydberg formula, find the wavelength for $i=1$ & $j=2,3,4,5$.

DE-BROGLIE'S EXPLANATION OF BOHR'S 2nd POSTULATE OF QUANTISATION

$$L_n = n \frac{h}{2\pi}$$

$2\pi r_n = n\lambda$ (Here, λ is the de-Broglie wavelength of electron)

$$\therefore 2\pi r_n = \frac{nh}{mv_n} \quad [\because \lambda = \frac{h}{p} \text{ or } \frac{h}{mv}]$$

$$\Rightarrow mv_n r_n = n \frac{h}{2\pi} \quad \text{OR} \quad L_n = \frac{nh}{2\pi}$$

Atomic Spectra

$$\frac{1}{\lambda} = R \left(\frac{1}{n_1^2} - \frac{1}{n_2^2} \right)$$

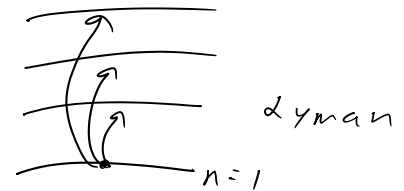
If $n_1 = 1 \Rightarrow$ Lyman Series (U.V.)

$n_1 = 2 \Rightarrow$ Balmer Series (Visible)

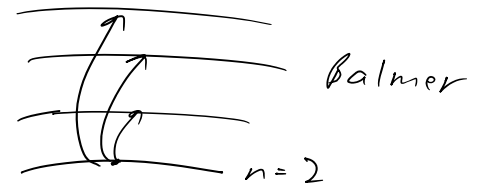
$n_1 = 3 \Rightarrow$ Paschen Series (Visible)

$n_1 = 4 \Rightarrow$ Brackett Series (I.R.)

$n_1 = 5 \Rightarrow$ Pfund Series (I.R.)



⊕



Eg:- Find the wavelengths of 1st 4 spectral lines of the Paschen Series.

H.W. :- Exercise (Ch-12)

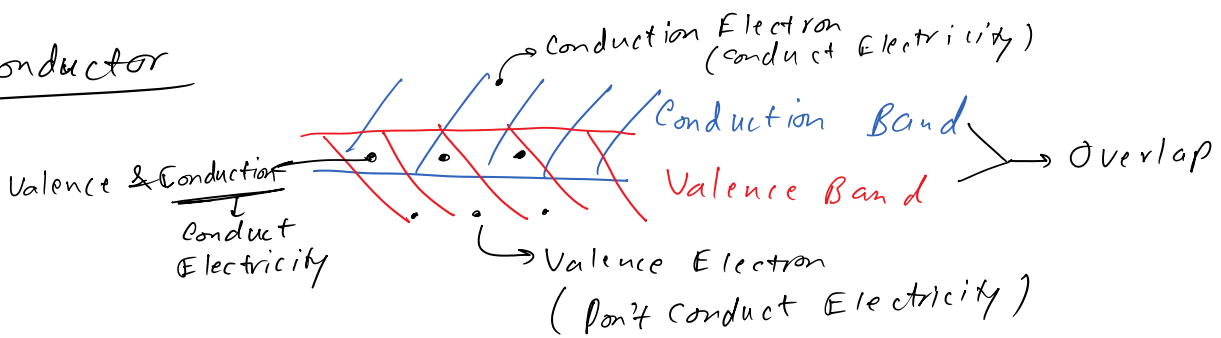
Energy Bands

(i) metals $\rightarrow \rho \sim 10^{-2} - 10^{-8} \Omega m^{-1}$ $\rho \rightarrow$ Resistivity
 $\sigma \sim 10^2 - 10^8 S m^{-1}$ $\sigma \rightarrow$ Conductivity

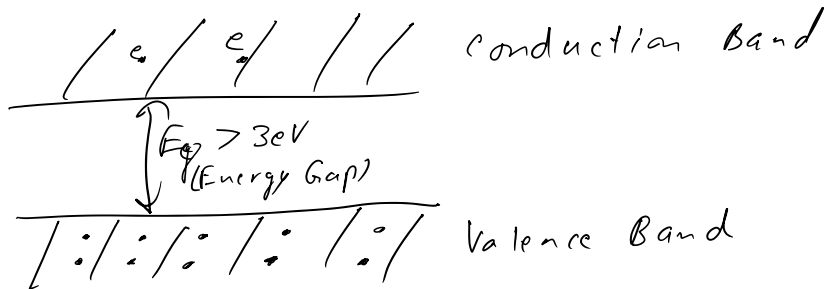
(ii) Semiconductors $\rightarrow \rho \sim 10^{-5} - 10^{-6} \Omega m^{-1}$
 $\sigma \sim 10^5 - 10^6 S m^{-1}$

(iii) Insulators $\rightarrow \rho \sim 10^{11} - 10^{19} \Omega m^{-1}$
 $\sigma \sim 10^{-11} - 10^{-19} S m^{-1}$

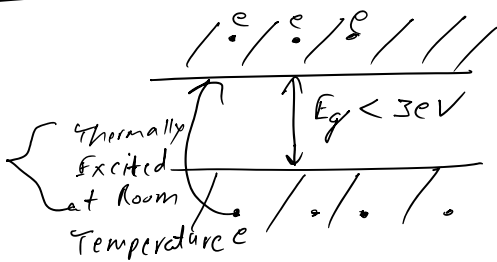
* Conductor



* Insulator

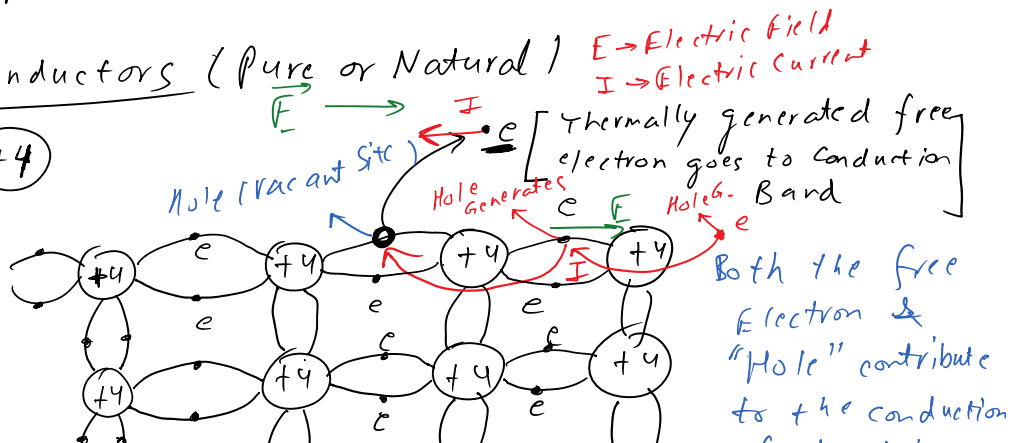


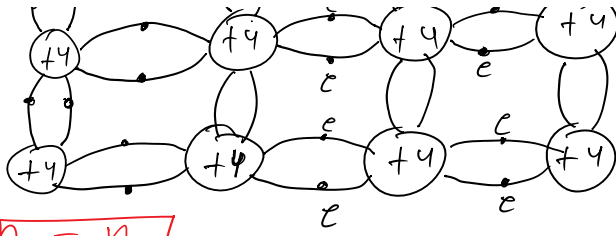
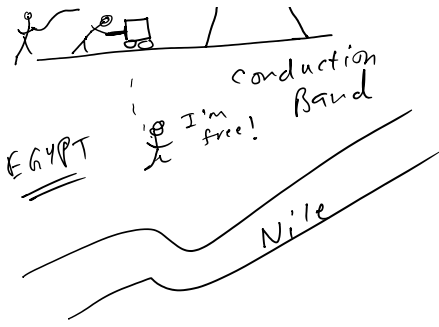
* Semiconductor



Intrinsic Semiconductors (Pure or Natural)

(Si or Ge) \rightarrow (+4)

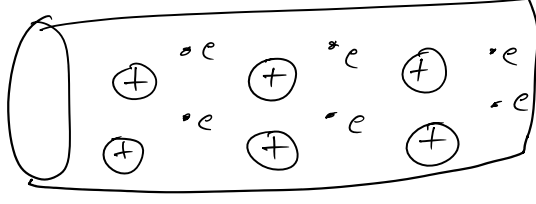




"Hole" contribute to the conduction of electricity.

$$n_e = n_h$$

In semiconductor, both the movement of free electron and hole leads to electric current.



In conductor, most electrons are free. In semiconductors, few electrons are free.

EXTRINSIC SEMICONDUCTOR (Doped)

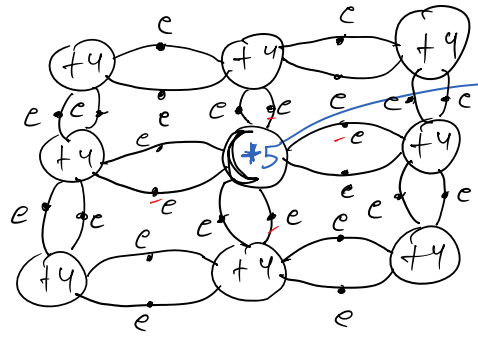
↳ Performance enhanced semiconductor

n-type

$$n_e \gg n_h \text{ intrinsic}$$

$$n_i^2 = n_e n_h$$

$n_i \rightarrow$ Intrinsic (Holes & Electrons)
 $n_e, n_h \rightarrow$ Extrinsic Carriers.



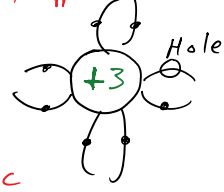
one extra electron per atom of the dopant is always free

* A large number of free electrons is generated by doping.

p-type

$$n_h \gg n_e$$

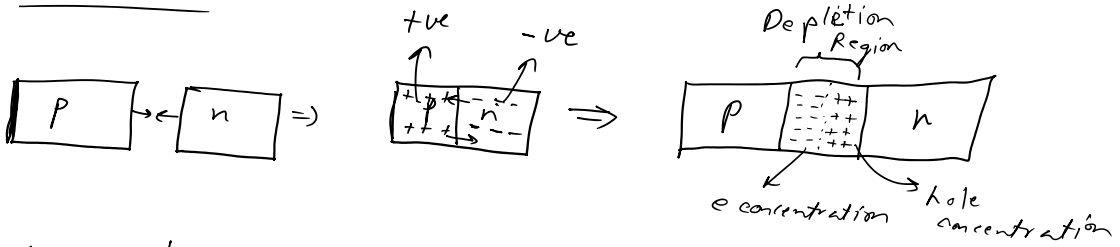
↳ intrinsic



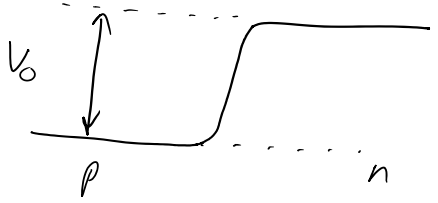
* Pentavalent Dopants like Arsenic (As), Antimony (Sb), Phosphorus (P).

* Trivalent Dopants like Al, Indium (In), Boron (B).

* P-n Junction

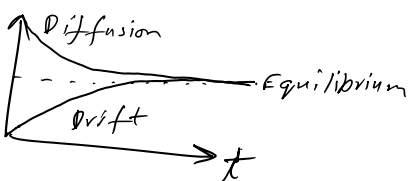
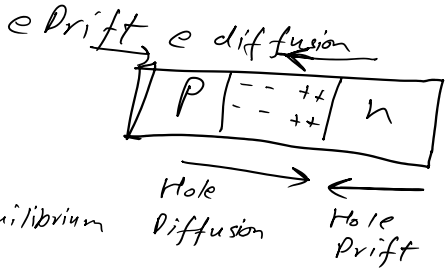
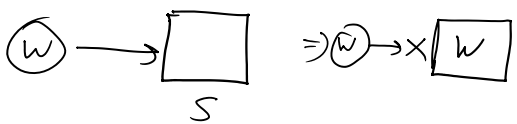


for an electron or hole to cross the depletion region, it requires energy to overcome a potential Barrier.

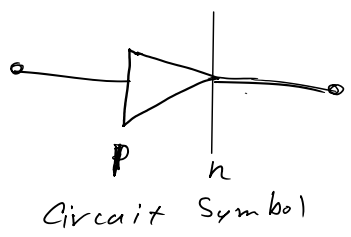
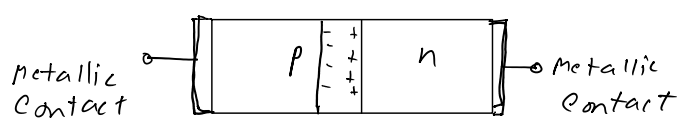


* There are two types of movements of the electrons & holes:
 (i) Drift (ii) Diffusion.
 * At the equilibrium stage of a

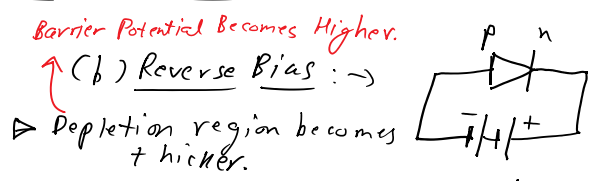
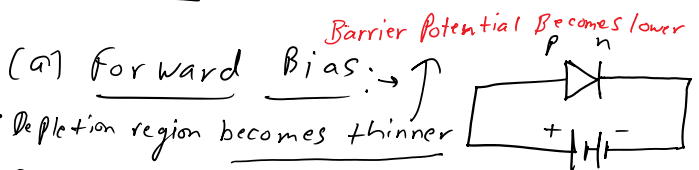
p n * At the equilibrium stage of a p-n junction, drift equals diffusion.



Semiconductor Diode (p-n junction diode)

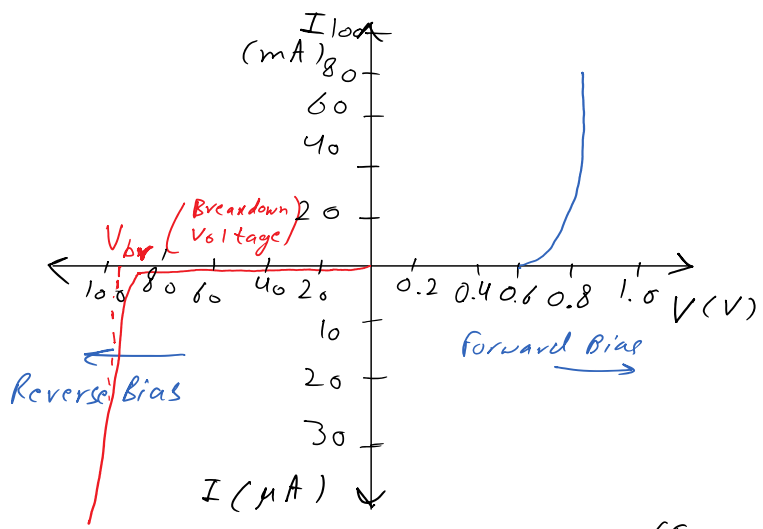


* There are two ways of connecting a diode to a battery



For an ideal diode, there is zero resistance in forward bias & infinite resistance is offered in reverse bias.

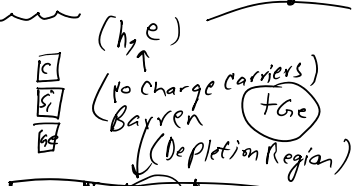
V-I characteristics of a diode



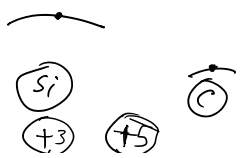
Eq 14.4

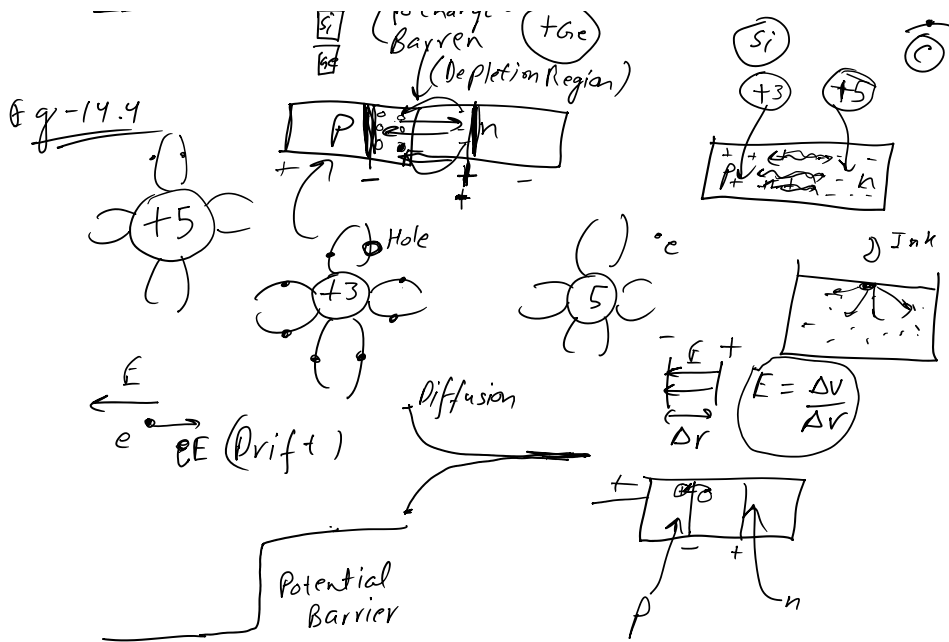
Ex-14.1 to 14.5

Ex-14.3



$(E_g)_c > (E_g)_s > (E_g)_{Ge}$

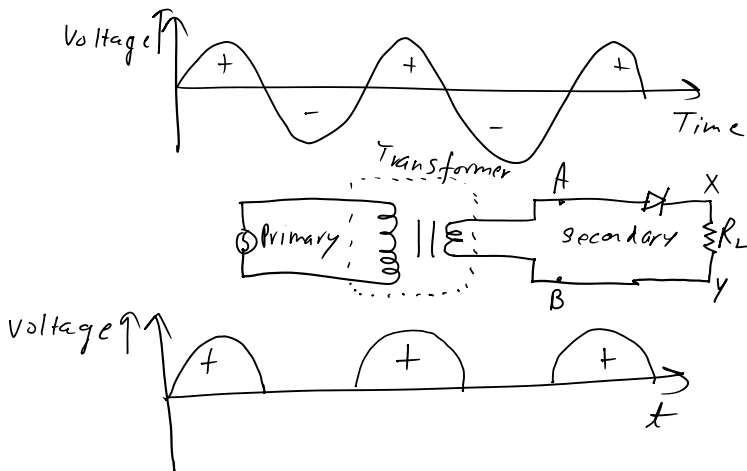




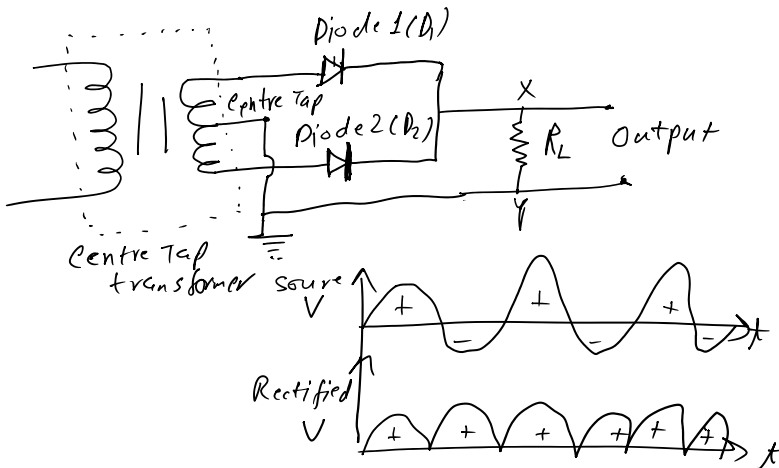
Application of Junction Diode as a Rectifier

↳ Converts AC to DC
Current

Half-Wave Rectifier



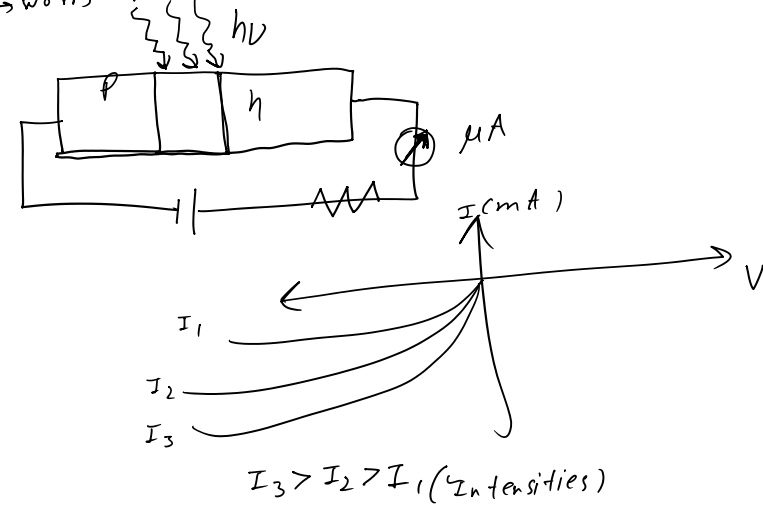
Full Wave Rectifier



Photodiode

↳ To measure the intensity of light. (Light Intensity Sensor)

Photo Diode
 To measure the intensity of light. (Light Intensity Sensor)
 works in reverse bias.



Light Emitting LED

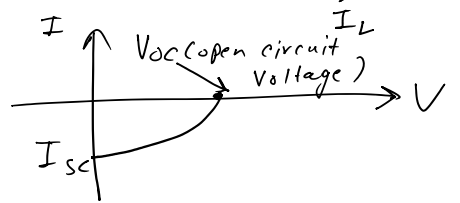
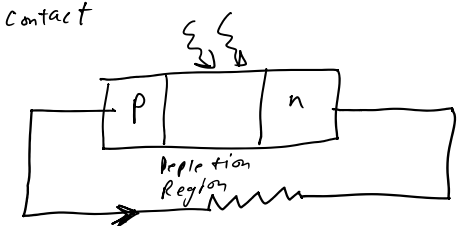
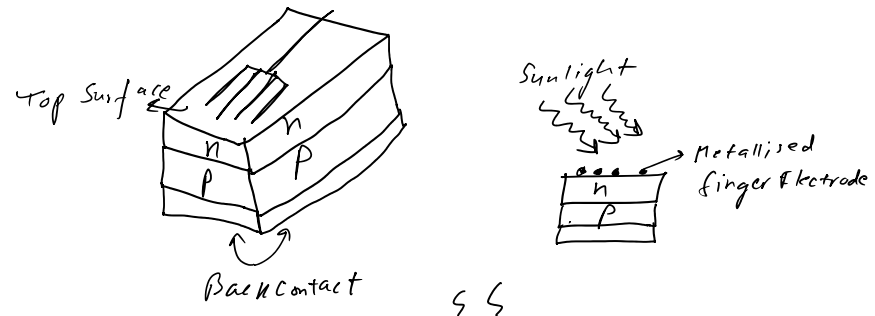
↳ P-n junction diode that emits light in forward bias.

Benefits of LED

- (i) Low operational voltage & less power
- (ii) Fast action and no warm up time required (unlike sodium lamps)
- (iii) The bandwidth of emitted light is 100 \AA to 500 \AA or in other words it is nearly ~~to~~ monochromatic.
- (iv) Long life & ruggedness.
- (v) fast on-off switching capability.

Solar Cell

↳ Generates Voltage (Emf) when light falls on them.



Exercise (1 hour)

- ▷ Lens Formula → Practice
- ▷ Lens Maker's Formula
- ▷ Refraction at a curved Surface

LENS FORMULA

$$\frac{1}{f} = \frac{1}{v} - \frac{1}{u}$$

Q1:- If an object of height 20 cm is placed to the left of a convex lens at a distance of 30 cm from the lens. What would be the focal length of the lens if height of image is :
(a) 20 cm → 28.57 cm
(b) 10 cm → 27.27 cm

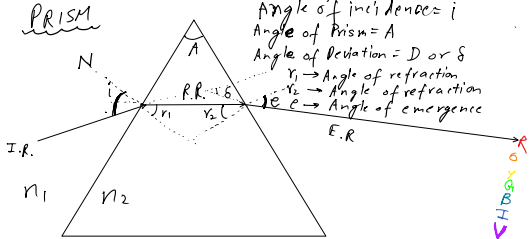
Magnification
 $m = \frac{v}{u} = \frac{h_2}{h_1}$

v = 90 cm
u = -45 cm

Q2:- Where should the object be placed in front of a convex lens of 130 cm if the image is twice (focal length) the size of the object.

Angle of incidence = i
Angle of Prism = A
Angle of Deviation = D or δ
r₁ → Angle of refraction
r₂ → Angle of refraction
e → Angle of emergence

n₁ → Refractive index of external medium
n₂ → Refractive index of the material of the prism.



$$\delta = i + e - A$$

$$n_2 = \frac{n_1 \sin \left[\frac{A + \delta_m}{2} \right]}{\sin \left[\frac{A}{2} \right]}$$

δ_m → Angle of minimum deviation.

Q1:- Find the angle of minimum deviation if angle of emergence is 37° and Angle of Prism is 4°.
Hint:- For minimum deviation i = e.

Sol:- $\delta = 37^\circ + 37^\circ - 4^\circ = 70^\circ$

$$\left(\frac{1.51}{1.33} \right) = \frac{(A + 6) / 2}{A / 2}$$

$$(1.14) = 1 + \frac{6}{A}$$

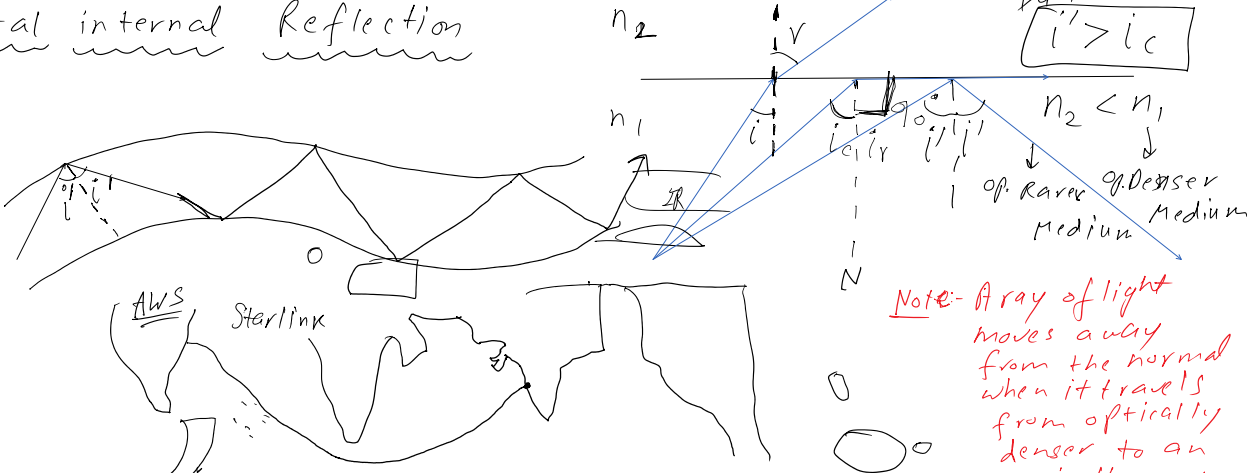
$$0.14 = \frac{6}{A} \Rightarrow A = \frac{6 \times 100}{14} = 42.85$$

Q2:- When a glass prism (n = 1.51) is placed in water (n = 1.33), the angle of minimum deviation is 6°. Find angle of Prism.

[Hint: When θ is small, sin θ ≈ θ]

$$A = 44^\circ$$

Total internal Reflection



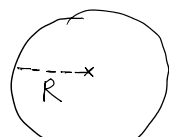
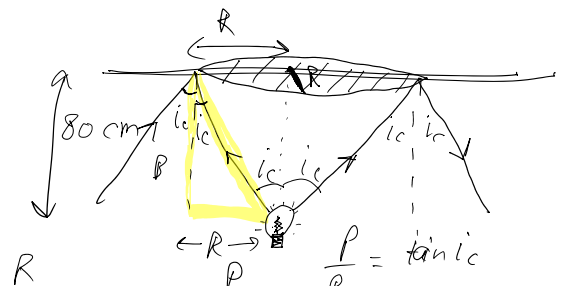
Note:- A ray of light moves away from the normal when it travels from optically denser to an optically rarer medium.

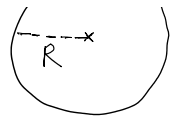
ic = Critical Angle
when (i = ic) r = 90°.

When i > ic, then i' = r (i')

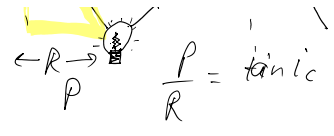
$$\sin ic = \frac{n_2}{n_1} = n_{21}$$

Ex - 9.5 :-
Pg 344
NCERT P-II





$$\frac{R}{10}$$

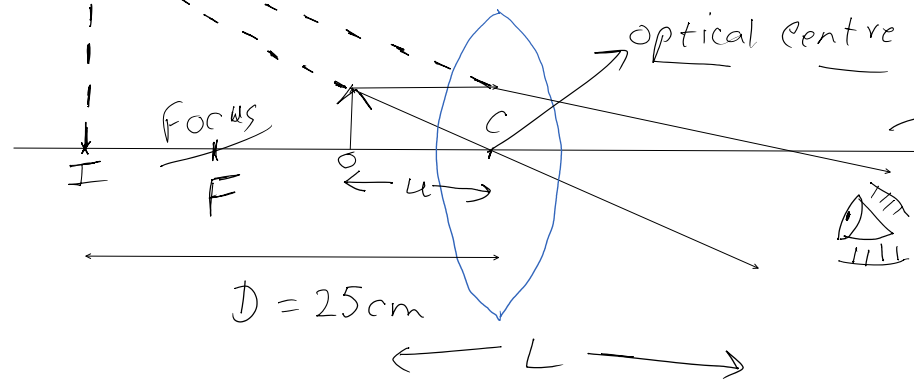


Optical Instruments

$$m = \frac{v}{u} = \frac{h_i}{h_o}$$

(i) Simple Microscope (Magnifying Glass)

Enlarged Virtual Image (Erect)

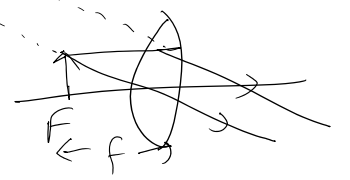


$$m_{\infty} = \frac{D}{f}$$

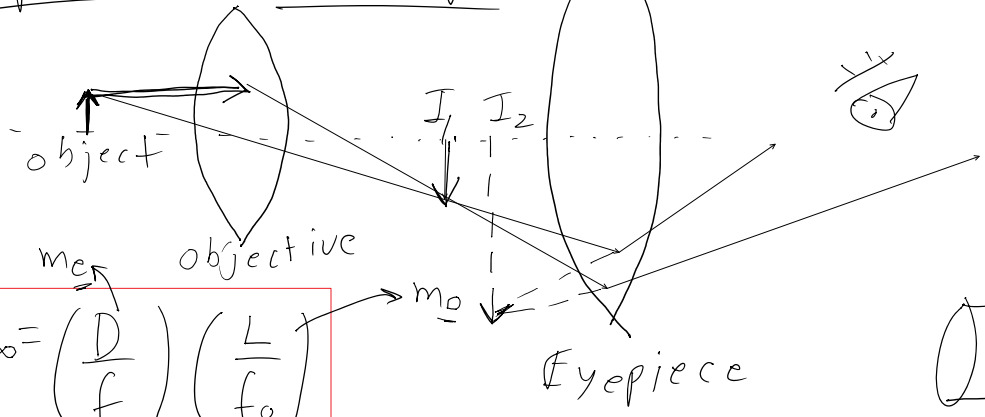
Image at infinity

$$m_D = 1 + \frac{D}{f}$$

Image at Near Point



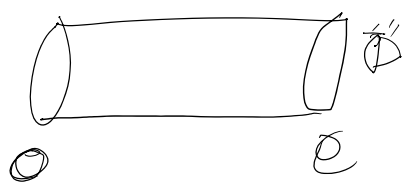
(ii) Compound Microscope



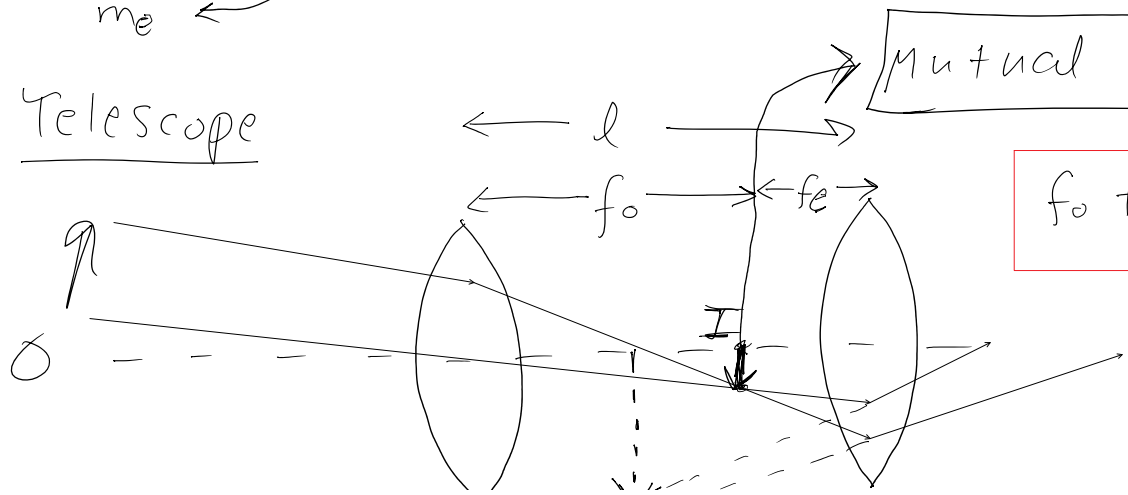
$$m = m_1 m_2$$

$$m_{\infty} = \left(\frac{D}{f_o}\right) \left(\frac{L}{f_e}\right)$$

$$m_D = \left(1 + \frac{D}{f_e}\right) \left(\frac{L}{f_o}\right)$$



Telescope



Mutual Focus

$$f_o + f_e = l$$

$$m = \frac{f_o}{f_e}$$

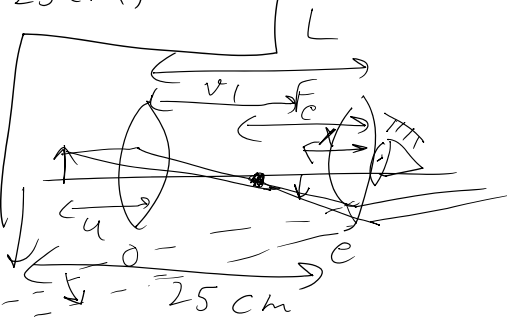
① Find the focal length of a magnifying glass of $m = 3$ with object placed at the focus.

formed
at
infinity
formed
at
point (D)
(25 cm)

comfortably
→ clear image.

$$\frac{25}{f} = 3$$

$$f = \frac{25}{3}$$

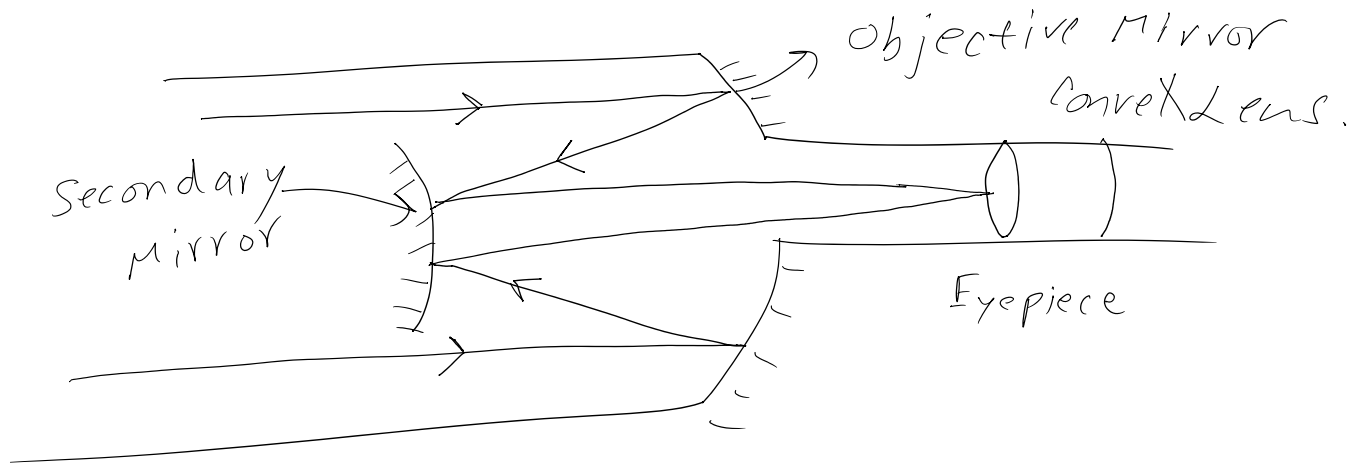


$m_1 m_2 \dots$



fe

Galilean Telescope OR Reflecting Telescope OR CASSEGRAIN TELESCOPE



* Cassegrain Telescope removes Chromatic Aberration (Multicolored Blurred Image)

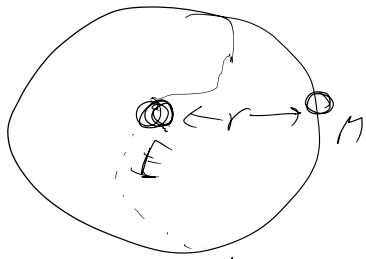
Exercise

9.5 A small bulb is placed at the bottom of a container containing water to a depth of t . What is the area of the surface through which light from the bulb can emerge? Refractive index of water is $\frac{4}{3}$. (Consider the bulb to be a point source)

Sol:- ??

SEGRAIN

→ Inventor



es)



om of a tank

80cm.

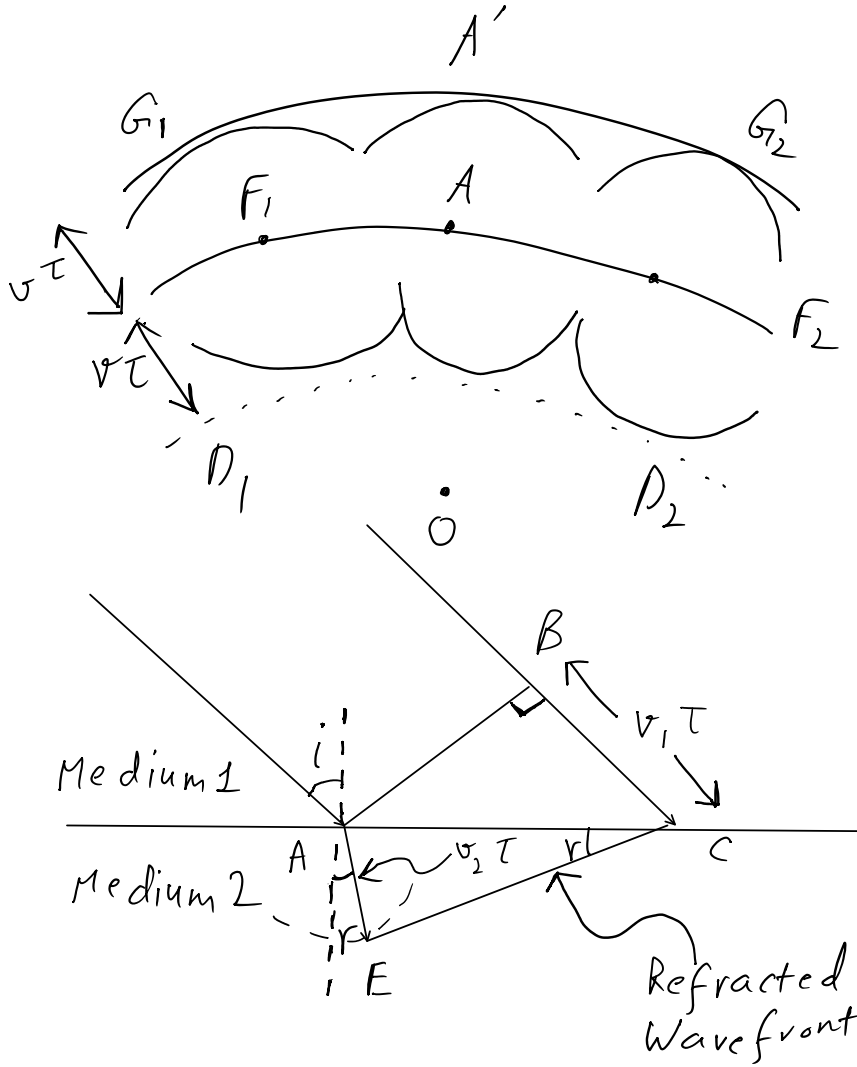
ace of water

16 can emerge

is 1.55.

source)

Huygen's Principle



$$\sin i = \frac{BC}{AC} = \frac{v_1 \tau}{AC}$$

$$\sin r = \frac{AE}{AC} = \frac{v_2 \tau}{AC}$$

$$\frac{\sin i}{\sin r} = \frac{v_1}{v_2}$$

$$n_1 = \frac{c}{v_1}$$

$$n_2 = \frac{c}{v_2}$$

$$\frac{\lambda_1}{\lambda_2} = \frac{BC}{AE} = \frac{v_1}{v_2}$$

$$\frac{v_1}{\lambda_1} = \frac{v_2}{\lambda_2}$$